1 The Celestial Sphere

Assign: Read Chapter 1 of Carrol and Ostlie (2006)

1.1 Geocentric Universe of the Greeks

1.2 Getting to Know the Celestial Sphere

- OJTA: 2. Overview of the Sky/The Celestial Sphere
	- **–** (1) The Celestial Sphere
	- **–** (2) The Ecliptic
	- **–** (3) The Coordinate System
	- **–** (4) Equinoxes and Solstices
	- **–** (5) Motion on the Celestial Sphere
- OJTA: 2. Overview of the Sky/Constellations
	- **–** (1) Groupings and Asterisms
	- **–** (2) Classical Constellations
	- **–** (3) Modern Constellations
	- **–** (6) Naming the Stars
- OJTA: 2. Overview of the Sky/Aspects and Phases
	- **–** (1) Classification
	- **–** (4) Wanderers
	- **–** Animation 3.2
- OJTA: 2. Overview of the Sky/Timekeeping
	- **–** (1) Sidereal and Solar Time
	- **–** (2) Sidereal and Solar Days
	- **–** (3) Precession of the Earth's Axis
	- **–** Animation 2.13
	- **–** Animation 3.2

– (5) Inferior Planets

– (6) Superior Planets

Example Problem: Synodic and sidereal orbital period

For synodic period S and sidereal period P , $\ddot{}$

$$
\frac{1}{S} = \begin{cases} \frac{1}{P} - \frac{1}{P_{\oplus}} & \text{(Inferior)}\\ \frac{1}{P_{\oplus}} - \frac{1}{P} & \text{(Superior)} \end{cases}
$$

Generally, $P_{\oplus} = 365.26$ days and For Mercury, $P \simeq 88$ days. Thus

$$
\frac{1}{S} = \frac{1}{88 \text{ d}} - \frac{1}{365.26 \text{ d}} = 8.63 \times 10^{-3} \text{ d}^{-1},
$$

so for Mercury $S = 115.9$ days.

- OJTA: 2. Overview of the Sky/The Seasons
	- **–** (1) The Northern Hemisphere
	- **–** (2) The Southern Hemisphere
- OJTA: 2. Overview of the Sky/The Moon (be brief)
	- **–** (1) Revolution in Orbit
	- **–** (2) Lunar Phases
	- **–** (3) Rotational Period
	- **–** (4) Tidal Locking (omit details til later)

1.3 Important "Rules of Thumb"

- Sun and Moon subtend about 1/2 degree (width of outstretched thumb)
- The sky appears to turn

 $\frac{360^{\circ}}{24 \text{ hr}} = 15^{\circ}/\text{hour}$ (Width outstretched pointer and pinkie) 360**^ı** $\frac{360}{24 \times 60} = 0.25^{\circ} / \text{ min}$ (Width outstretched thumb every 2 minutes)

This means that the sky turns about 1[°] every 4 minutes.

The Sun drifts eastward on the ecliptic

$$
\frac{360^{\circ}}{365.25} \simeq 1^{\circ} \text{ per day}
$$

The Moon drifts eastward with respect to the constellations

$$
\frac{360^{\circ}}{27.3} \simeq 13.2^{\circ} \text{ per day}
$$

 $(13.2/15) \times 60 \simeq 52.8$ minutes later rising per day

1.4 Angular Measure

• 360° in circle

$$
\bullet \left(\frac{1}{60}\right)^{\circ} = 1 \text{ arcminute } ()
$$

$$
\bullet \left(\frac{1}{60 \times 60}\right)^{\circ} = \left(\frac{1}{3600}\right)^{\circ} = 1 \text{ arcsecond } ()
$$

- 1 radian (rad) = $\frac{180}{\pi} \simeq 57.3^{\circ}$
- 1 arcsecond = 4.848×10^{-6} rad

One arcsecond is the angle subtended by a dime viewed at a distance of 2 km! Many properties in astronomy require measuring angles of this size or smaller.