# **1** The Celestial Sphere

Assign: Read Chapter 1 of Carrol and Ostlie (2006)

## **1.1** Geocentric Universe of the Greeks

#### **1.2** Getting to Know the Celestial Sphere

- OJTA: 2. Overview of the Sky/The Celestial Sphere
  - (1) The Celestial Sphere
  - (2) The Ecliptic
  - (3) The Coordinate System
  - (4) Equinoxes and Solstices
  - (5) Motion on the Celestial Sphere
- OJTA: 2. Overview of the Sky/Constellations
  - (1) Groupings and Asterisms
  - (2) Classical Constellations
  - (3) Modern Constellations
  - (6) Naming the Stars
- OJTA: 2. Overview of the Sky/Aspects and Phases
  - (1) Classification
  - (4) Wanderers
  - Animation 3.2
- OJTA: 2. Overview of the Sky/Timekeeping
  - (1) Sidereal and Solar Time
  - (2) Sidereal and Solar Days
  - (3) Precession of the Earth's Axis
  - Animation 2.13
  - Animation 3.2

- (5) Inferior Planets

- (6) Superior Planets

Example Problem: Synodic and sidereal orbital period

For synodic period S and sidereal period P,

$$\frac{1}{S} = \begin{cases} \frac{1}{P} - \frac{1}{P_{\oplus}} & \text{(Inferior)} \\ \frac{1}{P_{\oplus}} - \frac{1}{P} & \text{(Superior)} \end{cases}$$

Generally,  $P_{\oplus} = 365.26$  days and For Mercury,  $P \simeq 88$  days. Thus

$$\frac{1}{S} = \frac{1}{88 \text{ d}} - \frac{1}{365.26 \text{ d}} = 8.63 \times 10^{-3} \text{ d}^{-1},$$

so for Mercury S = 115.9 days.

- OJTA: 2. Overview of the Sky/The Seasons
  - (1) The Northern Hemisphere
  - (2) The Southern Hemisphere
- OJTA: 2. Overview of the Sky/The Moon (be brief)
  - (1) Revolution in Orbit
  - (2) Lunar Phases
  - (3) Rotational Period
  - (4) Tidal Locking (omit details til later)

### 1.3 Important "Rules of Thumb"

- Sun and Moon subtend about 1/2 degree (width of outstretched thumb)
- The sky appears to turn

 $\frac{360^{\circ}}{24 \text{ hr}} = 15^{\circ}/\text{hour} \quad \text{(Width outstretched pointer and pinkie)}$  $\frac{360^{\circ}}{24 \times 60} = 0.25^{\circ}/\text{ min} \quad \text{(Width outstretched thumb every 2 minutes)}$ 

This means that the sky turns about 1° every 4 minutes.

• The Sun drifts eastward on the ecliptic

$$\frac{360^{\circ}}{365.25} \simeq 1^{\circ} \text{ per day}$$

• The Moon drifts eastward with respect to the constellations

$$\frac{360^{\circ}}{27.3} \simeq 13.2^{\circ} \text{ per day}$$

 $(13.2/15) \times 60 \simeq 52.8$  minutes later rising per day

## 1.4 Angular Measure

• 360° in circle

• 
$$\left(\frac{1}{60}\right)^{\circ} = 1$$
 arcminute (')  
•  $\left(\frac{1}{60 \times 60}\right)^{\circ} = \left(\frac{1}{3600}\right)^{\circ} = 1$  arcsecond (")

- 1 radian (rad) =  $\frac{180}{\pi} \simeq 57.3^{\circ}$
- 1 arcsecond =  $4.848 \times 10^{-6}$  rad

One arcsecond is the angle subtended by a dime viewed at a distance of 2 km! Many properties in astronomy require measuring angles of this size or smaller.